

How to detect extrasolar planets? (= planets outside of our solar system)

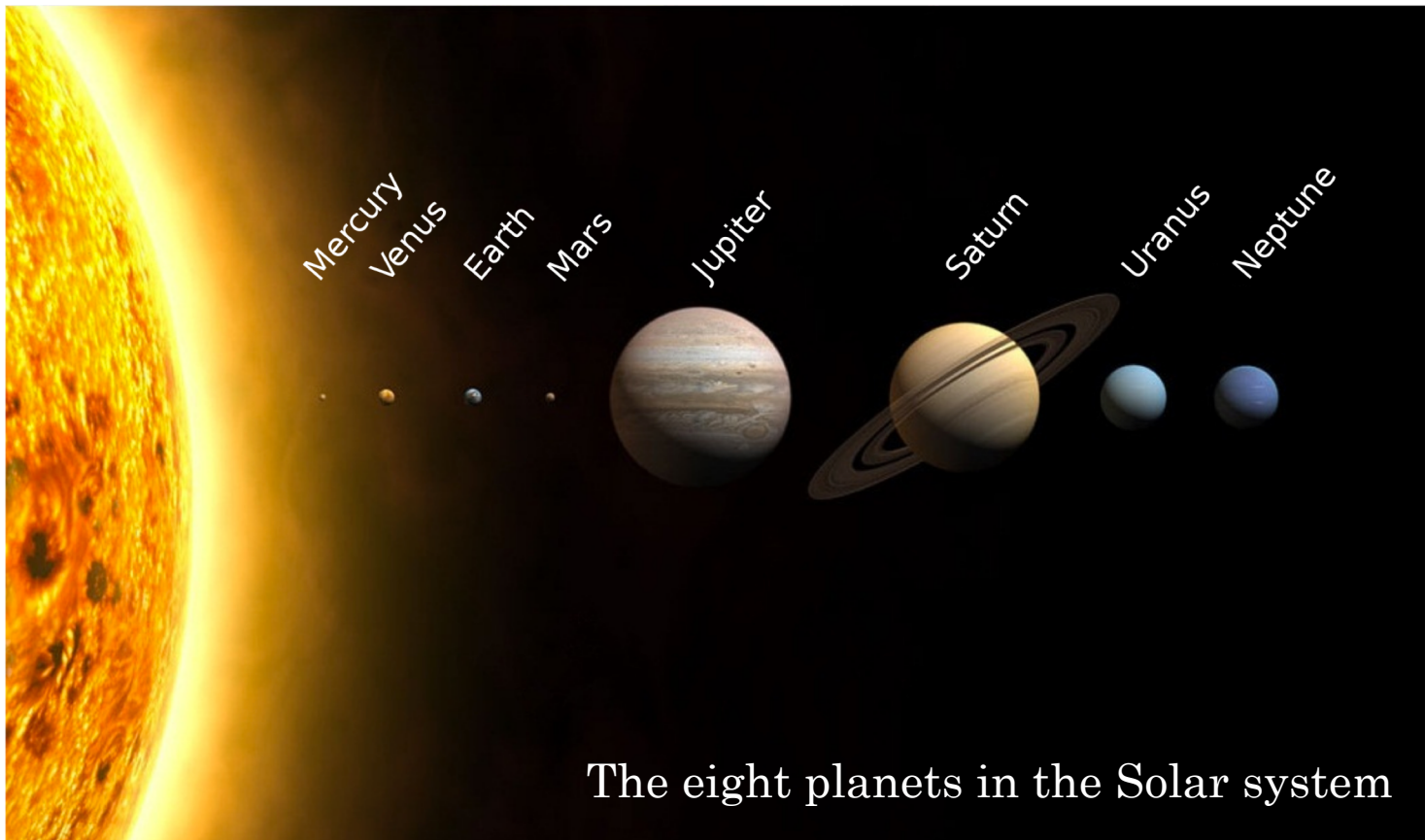
Dr. Denis Defrère

Inleiding tot de Sterrenkunde -- KUL (March 10th 2020)

What is a planet?

A planet is a celestial body which, in the Solar System,

- is in **orbit around the Sun**;
- has sufficient mass to assume **hydrostatic equilibrium** (a nearly round shape);
- has "**cleared the neighborhood**" around its orbit.



The eight planets in the Solar system

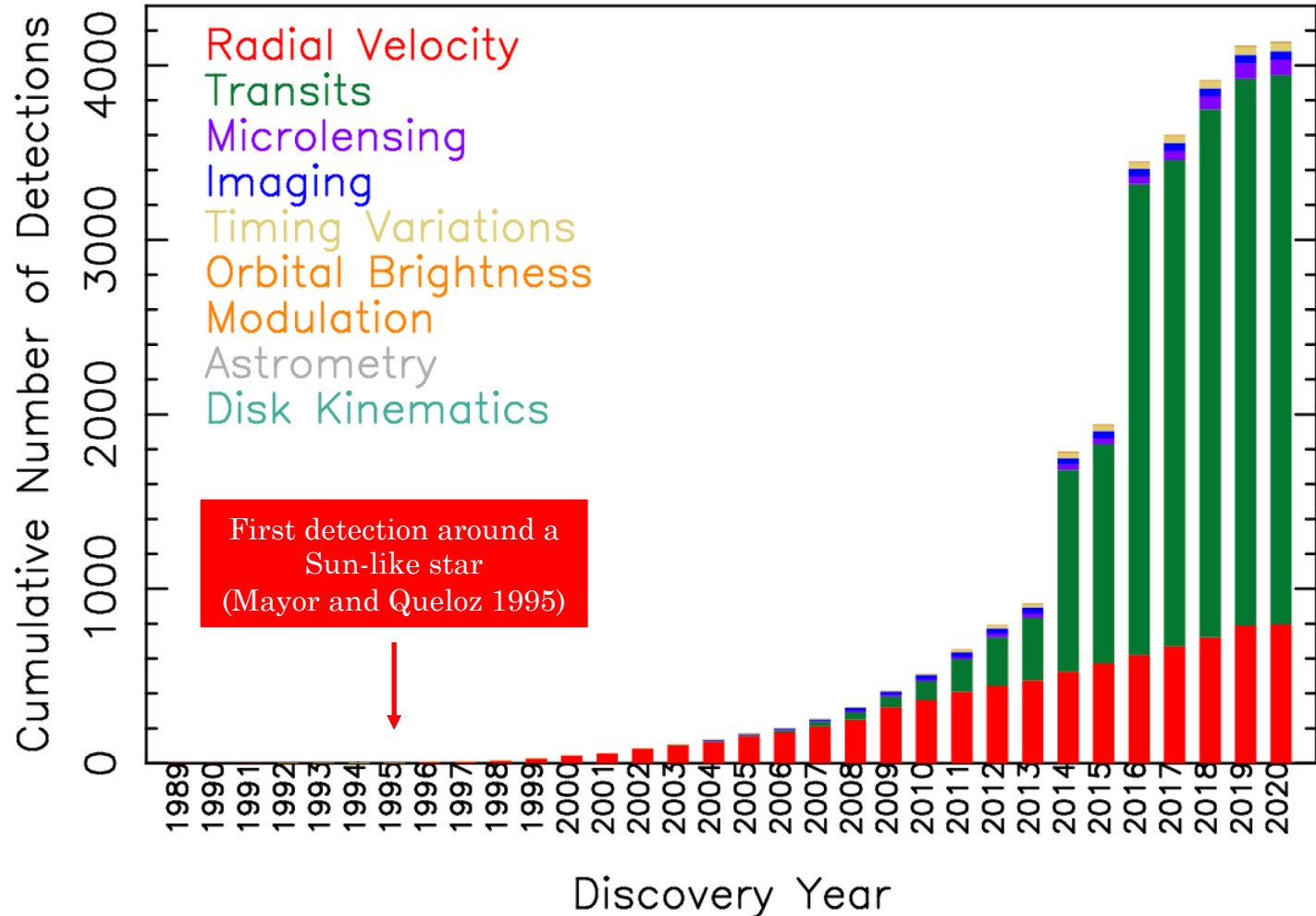
Questions:

Do you know **how many** extrasolar planets have been detected and **how**?

Lots of extrasolar planets!

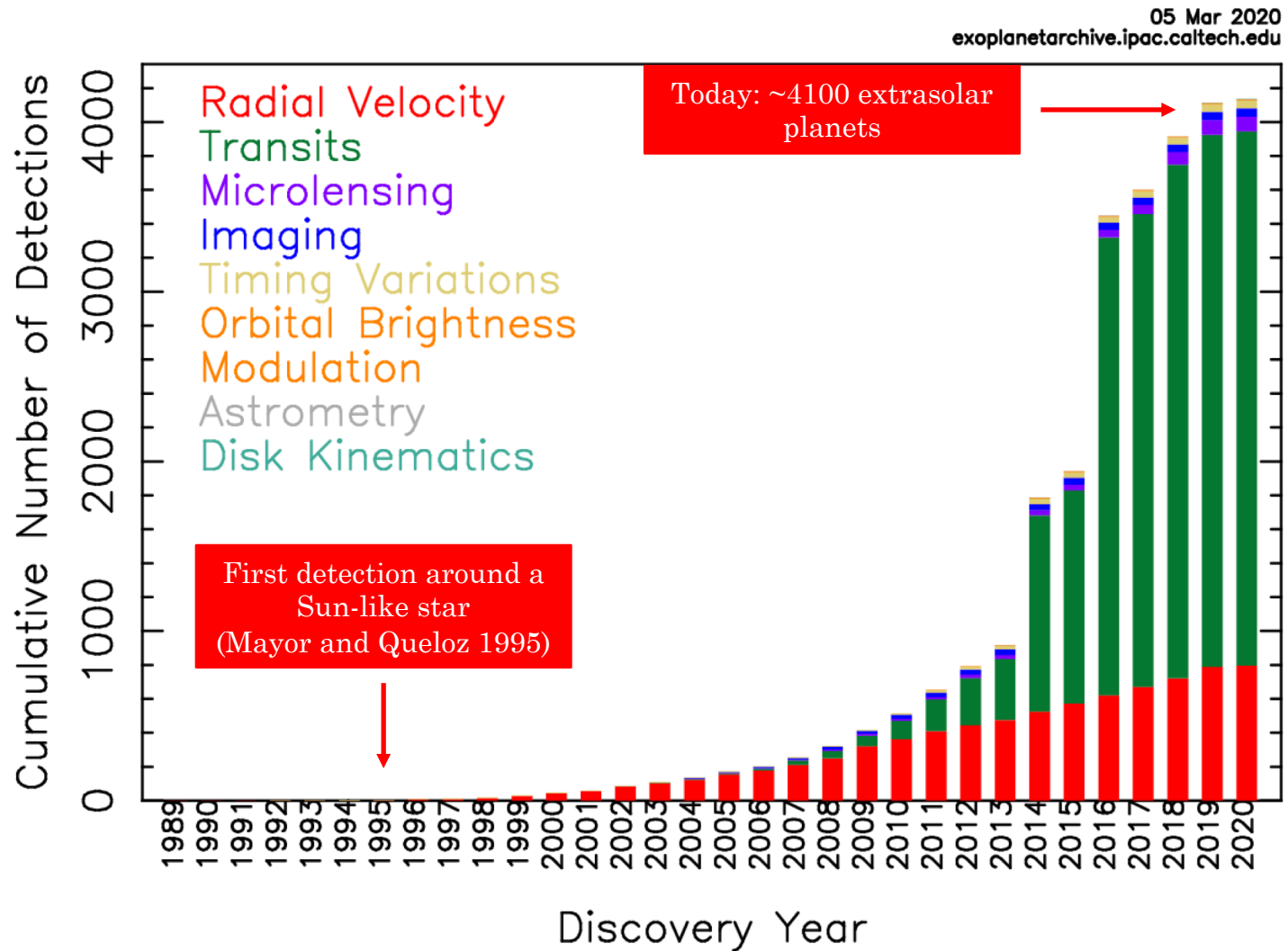
Cumulative Detections Per Year

05 Mar 2020
exoplanetarchive.ipac.caltech.edu



Lots of extrasolar planets!

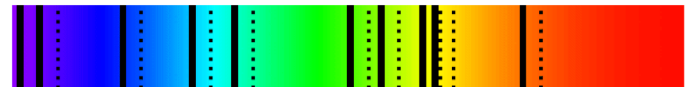
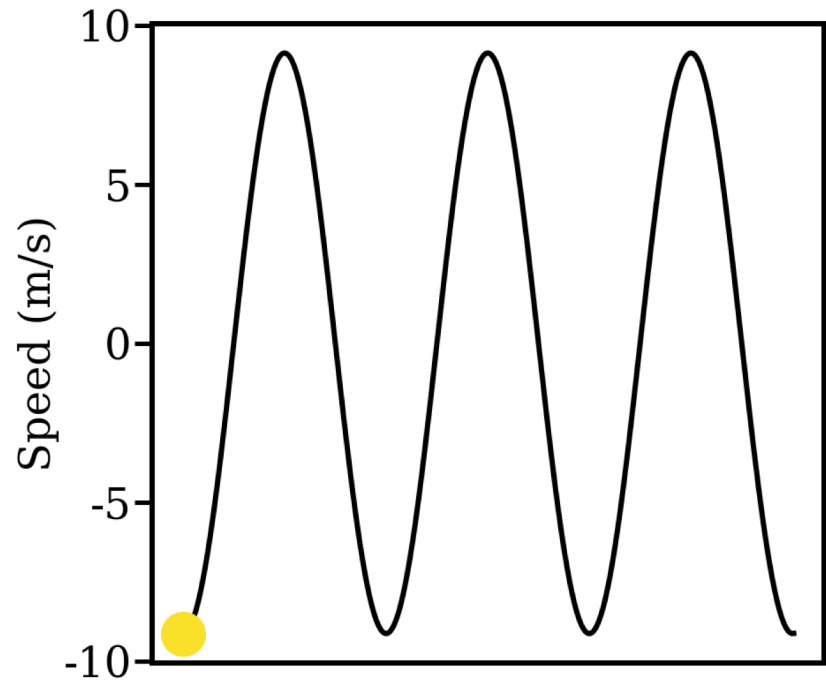
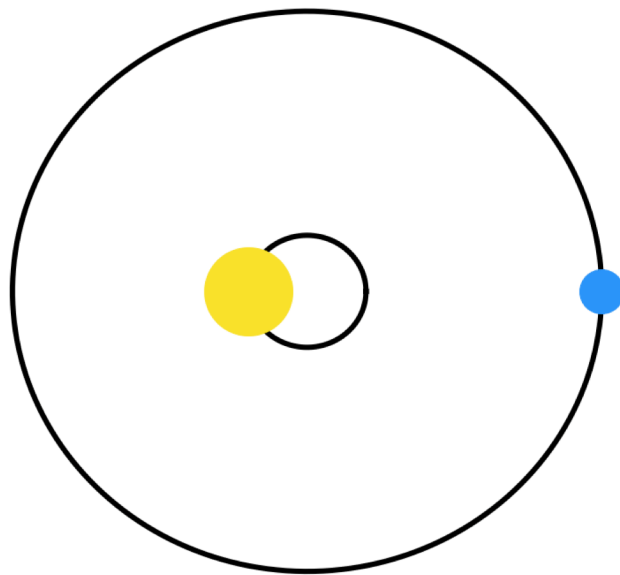
Cumulative Detections Per Year



Technique 1: Radial velocity

- Star, planet move around common center of mass
- Doppler effect moves spectral lines
- Look for periodic variations in stellar velocity

Alysa Obertas (@AstroAlysa)



Technique 1: RV signal

- Semi-amplitude of radial velocity given by:

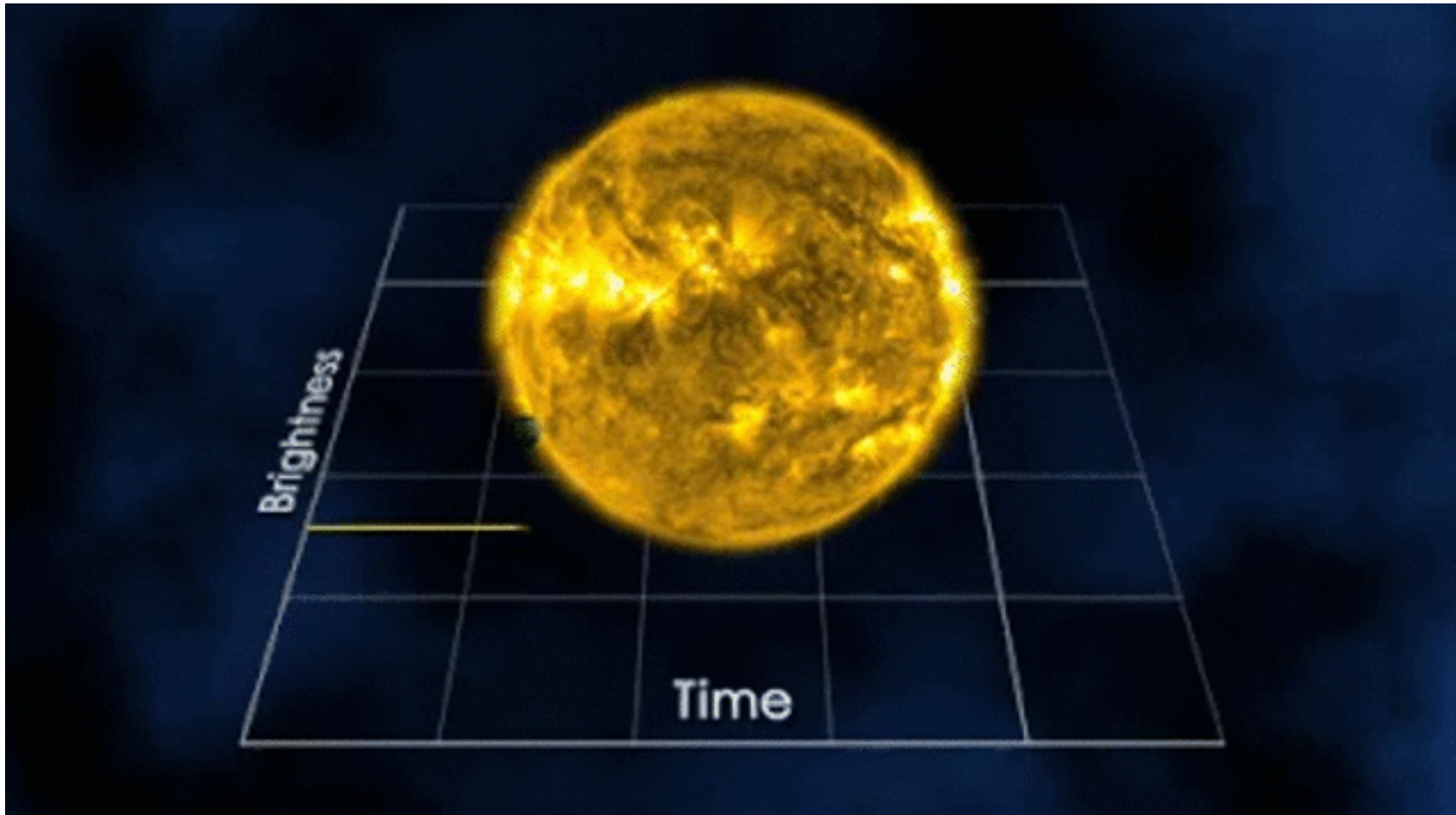
$$K = \left(\frac{2\pi G}{P_{orb}} \right)^{1/3} \frac{M_p \sin i}{(M_* + M_p)} \frac{1}{\sqrt{1 - e^2}}$$

*measured
*derived

- P_{orb} : orbital period
- M_* : mass of star
- M_p : mass of planet
- i : inclination, angle between normal to orbital plane and line of sight
- e : eccentricity

Technique 2: transit

- Low probability but simple observation
- Good for large planets close to the star

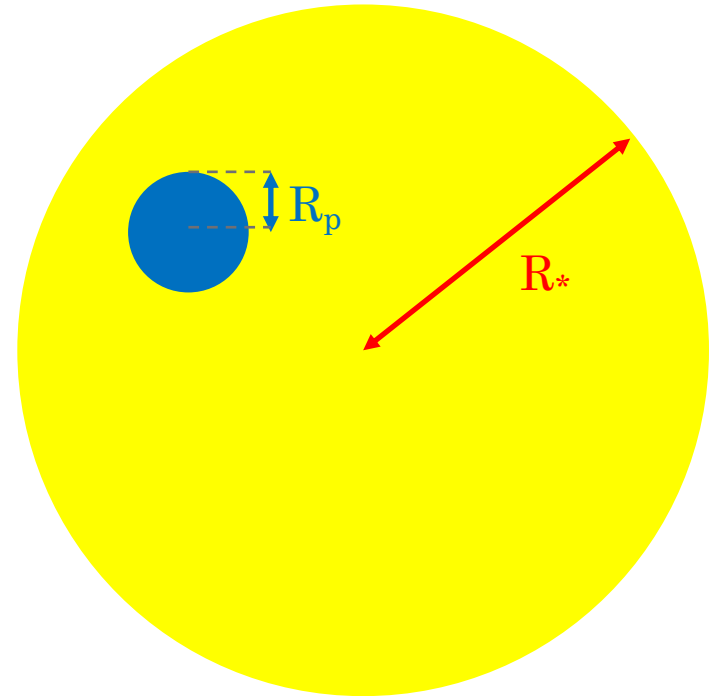


Technique 2: transit signal

- Intensity signal:

$$\frac{\Delta I}{I} = \left(\frac{R_p}{R_*} \right)^2$$

*measured
*derived

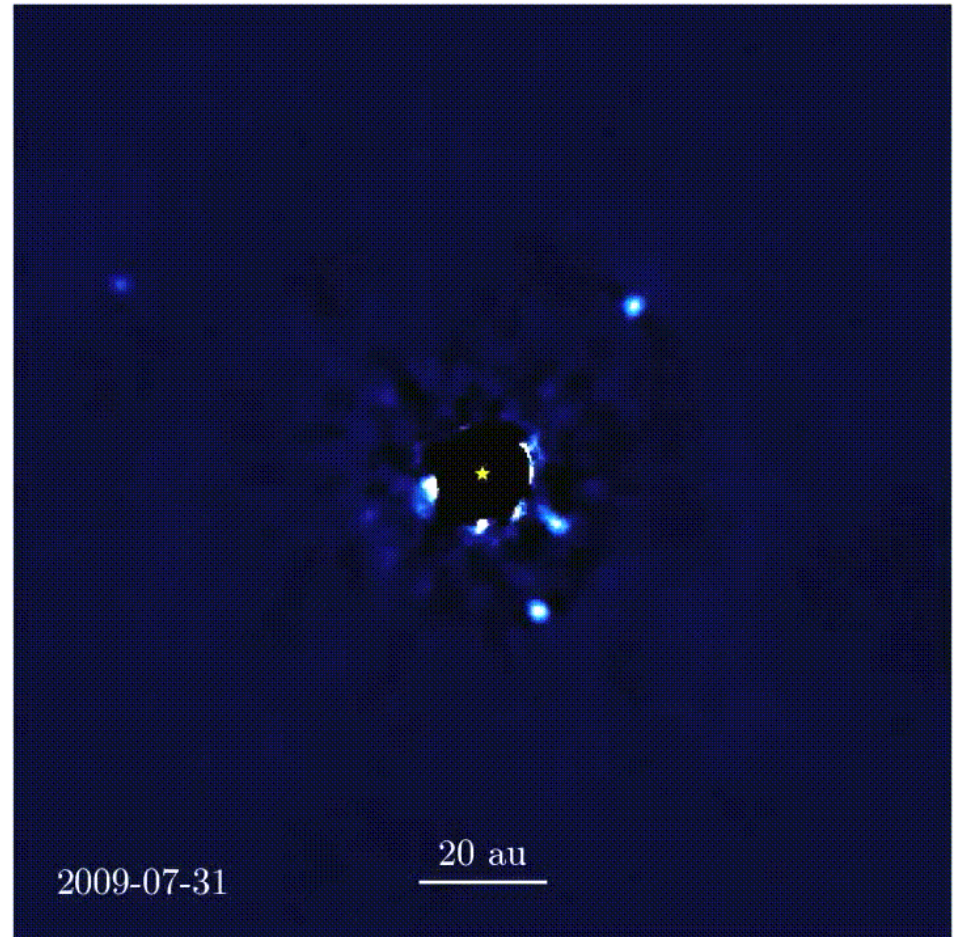


- R_* : stellar radius
- R_p : planet radius
- About 1% for Jupiter and Sun
- Transit duration proportional to $(P_{orb})^{1/3} R_* / (M_*)^{1/3}$
- Transit duration: also estimate of stellar radius
- Intensity change then provides planetary radius

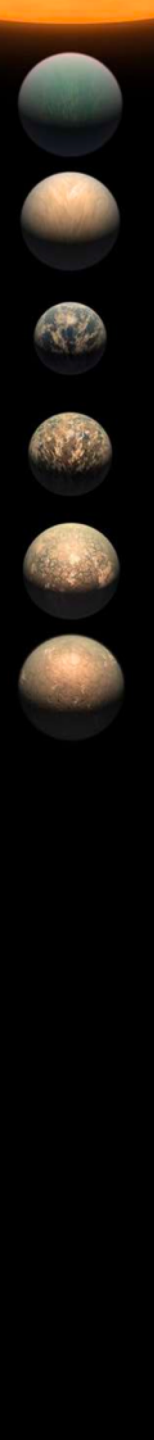
Technique 3: direct imaging

- Equivalent to taking a **picture** of the planetary system
- Require good sensitivity and high angular resolution => large telescope

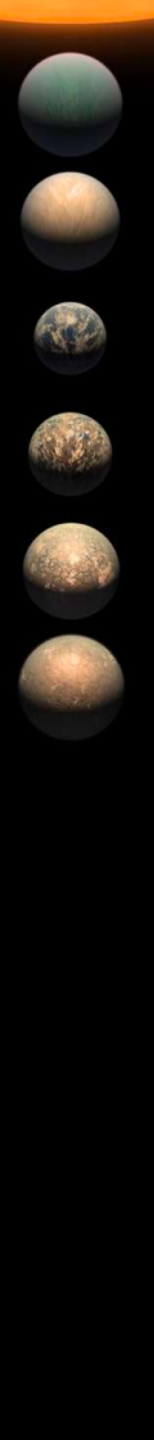
HR 8799 surrounded
by 4 giants extrasolar
planets



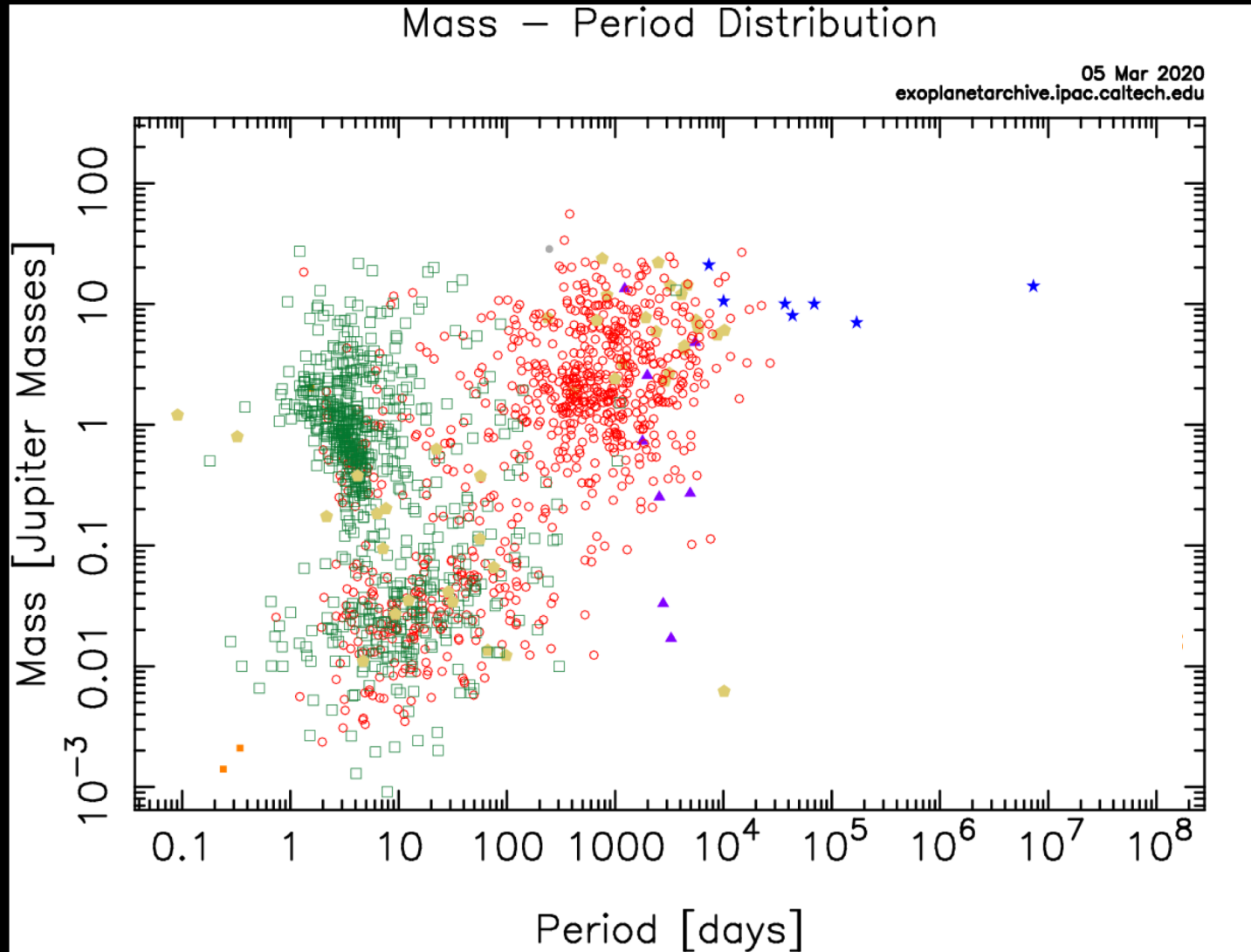
Technique 3: direct imaging



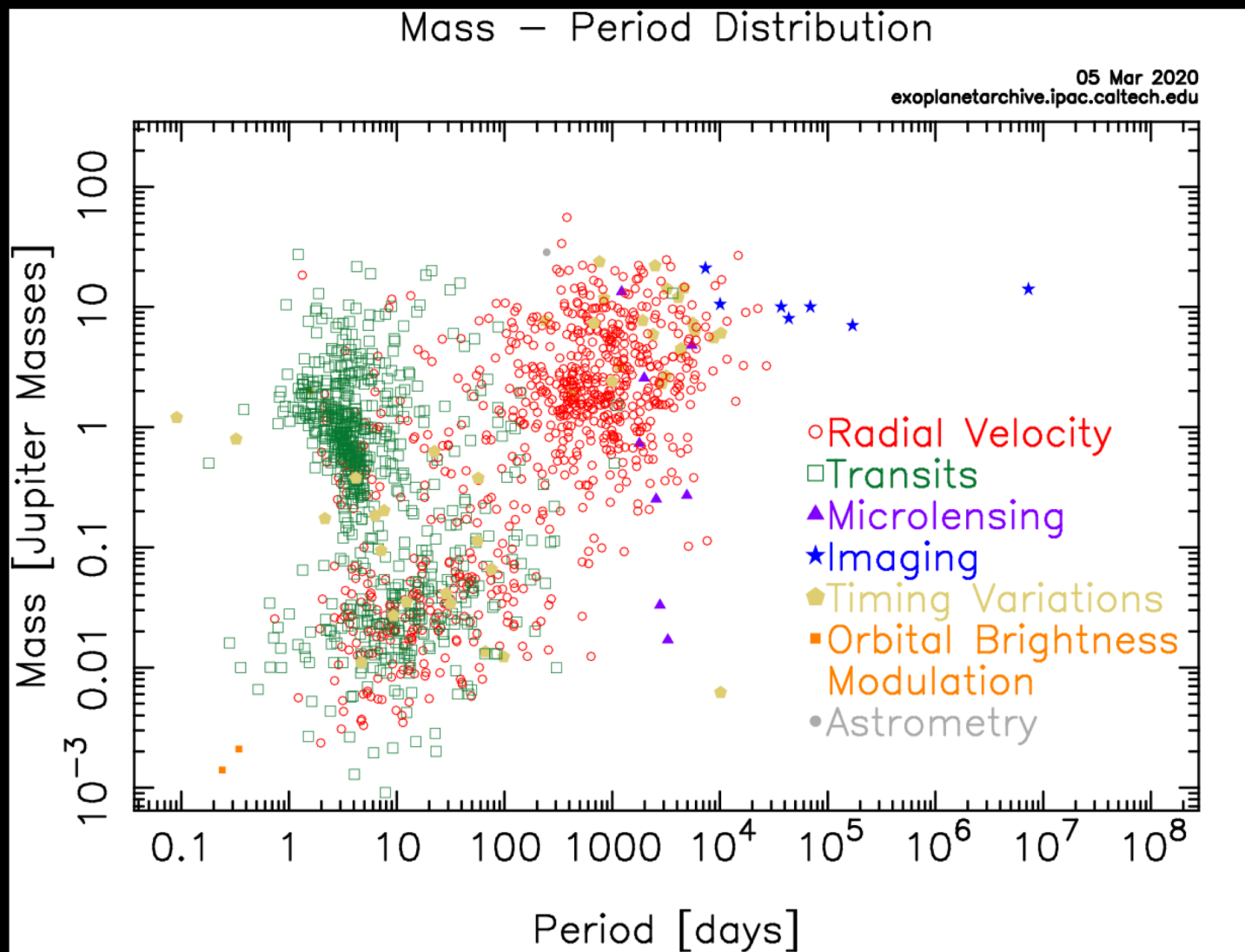
Technique 3: direct imaging



Questions: What kind of extrasolar planets can we detect today?



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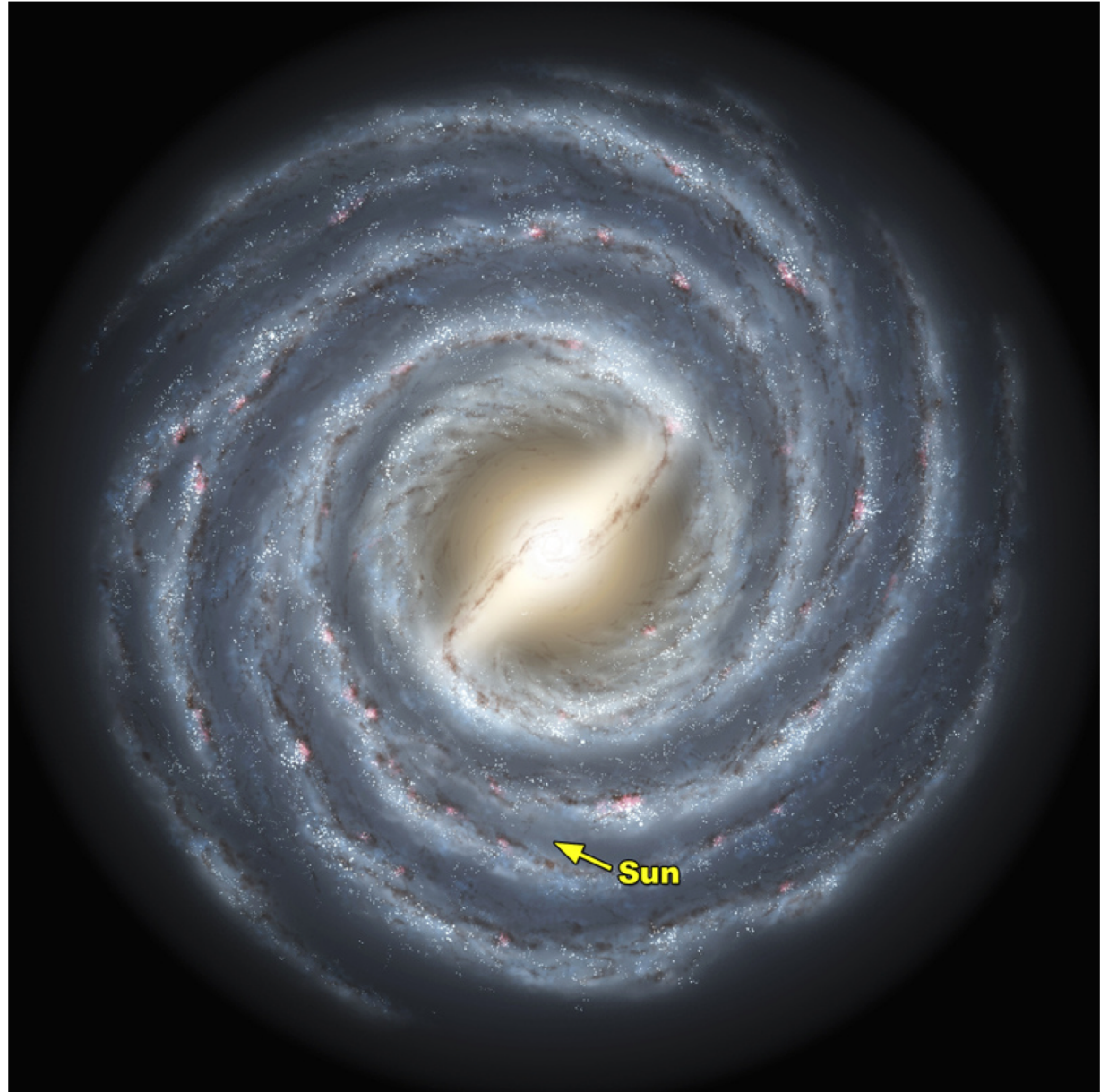


What we have learned today

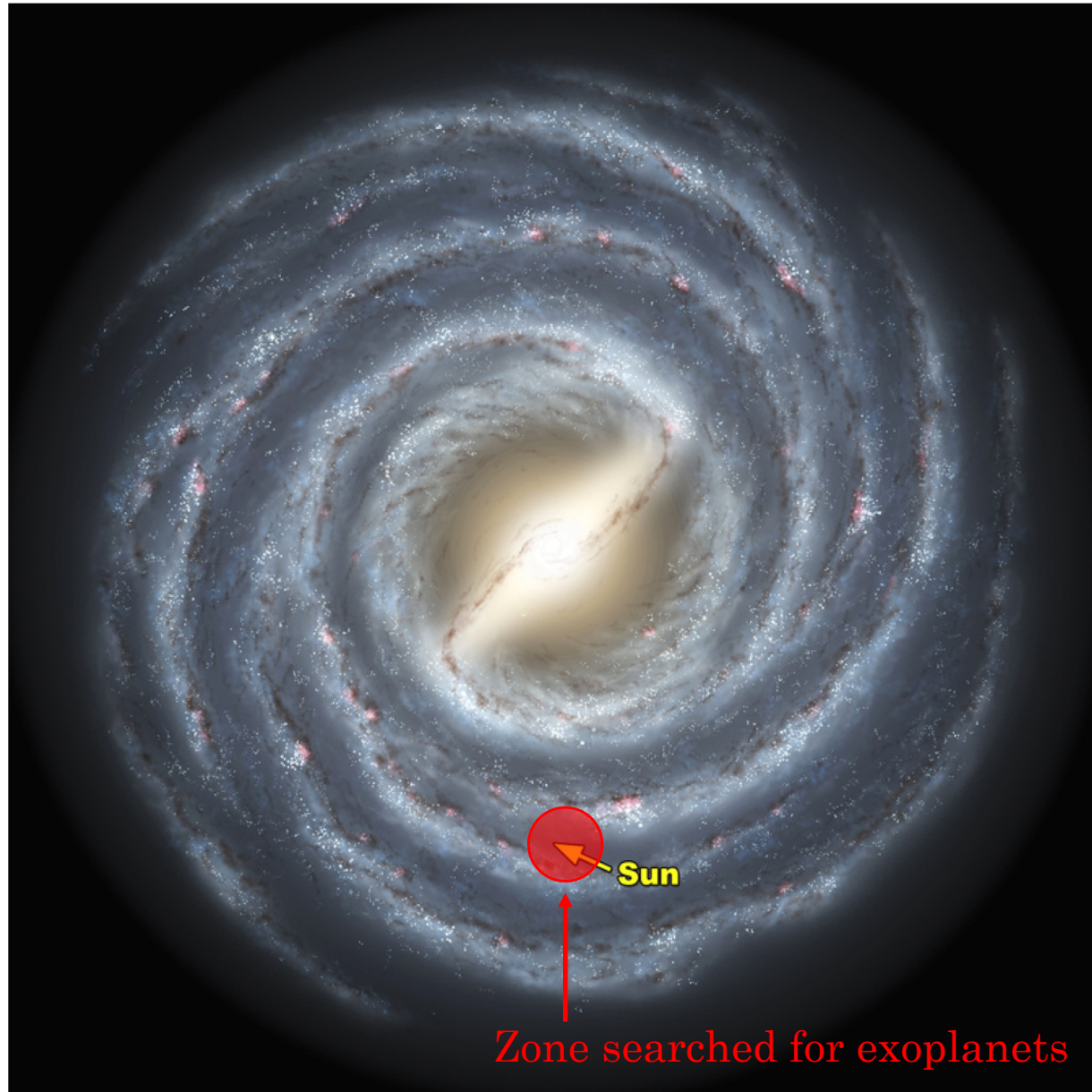
- First extrasolar planet around a Sun-like star detected in 1995
- More than **4100 extrasolar** planets detected today
- Different detection techniques (radial velocity, transit, direct imaging, ...)
- What we can learn: presence, mass, radius, orbit
- Next: how to characterize them? (i.e. chemical composition)



Some perspectives



Some perspectives



Further readings

- A survey of exoplanet detection techniques:
<https://arxiv.org/pdf/1805.02771.pdf>
- NASA exoplanet archive for up-to-date plots and latest results:
<https://exoplanetarchive.ipac.caltech.edu/>
- Extrasolar planet encyclopedia for news and catalog: <http://exoplanet.eu/>
- Slides of this presentation:
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